STATUS OF LEVEL 2 PROCESSING

AIRS SCIENCE TEAM MEETING

SEPTEMBER 18-20, 2002

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OUTLINE

This is a Preliminary Status Report

Caveats

Radiance biases vs. ECMWF for clear ocean cases

IR Tuning Methodology

Effect of IR tuning on sounding accuracy for clear cases

RMS sounding errors for cloudy cases

Biases of clear column radiances vs. ECMWF for cloudy ocean

Comparison of Total O₃ with TOMS data

CAVEATS

All calculations use pre-launch transmittances -

Frequencies are incorrect

Physics has been improved

"Tuning" to account for systematic errors is necessary

IR channels are treated as "noisy" because of residual computational errors

IR channels receive less weight in the solution

4.3 µm channels currently not used

Effects of non-LTE during day

Retrieval biases increase at night - current tuning insufficient

No angle correction is applied

No first product step is done

Microwave product is used as initial guess

CAVEATS (cont.)

"Errors" are differences from ECMWF forecast

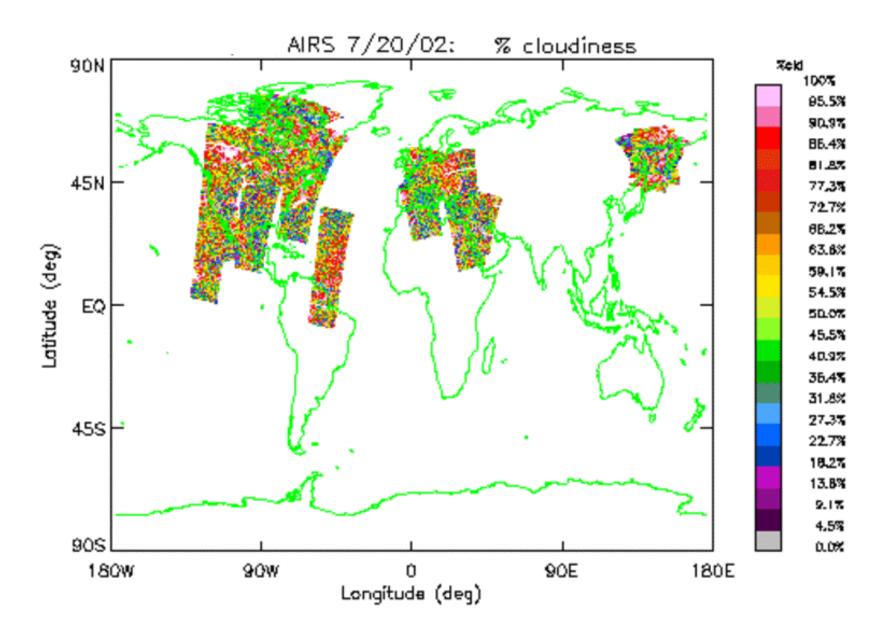
Actual errors are less

Forecast is not perfect

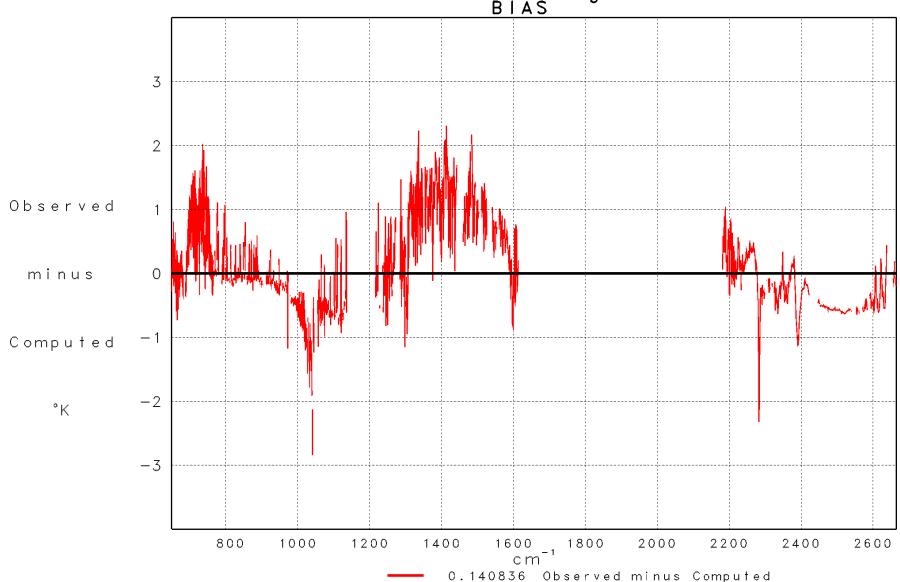
Forecast is every 3 hours - \pm 1 1/2 hours from observations

Spatial structure of forecast is coarser than that of real atmosphere

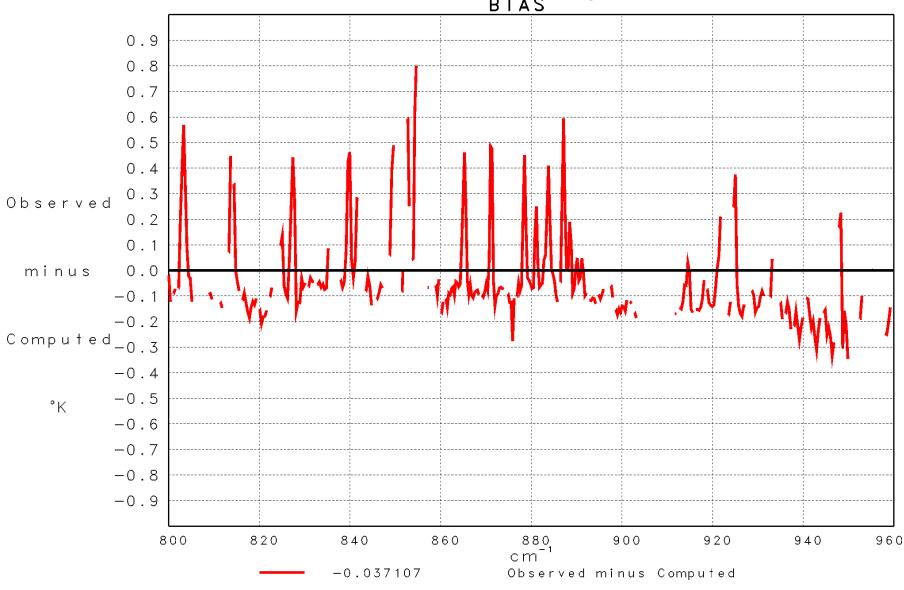
Errors should eventually be much smaller



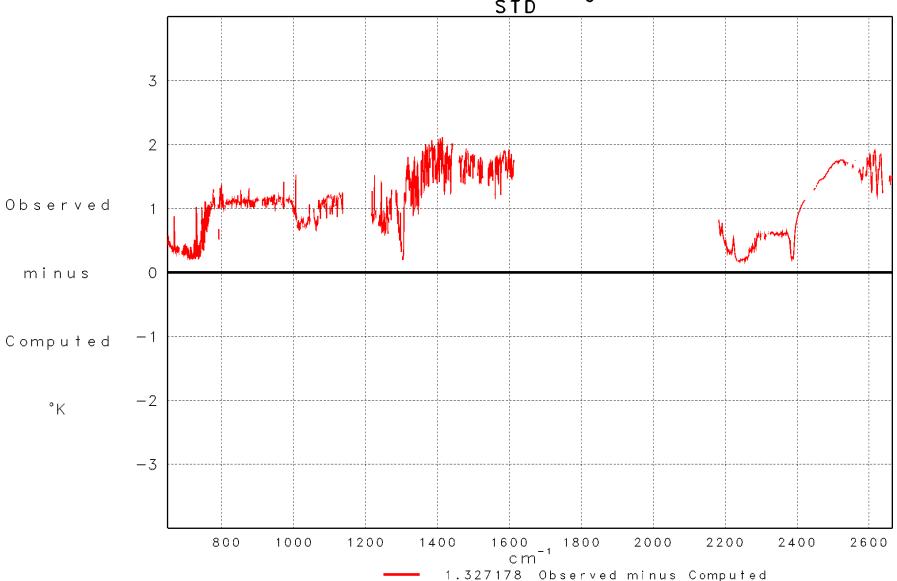
Brightness Temperature Difference 466 Clear Ocean Night Cases BIAS



Brightness Temperature Difference 466 Clear Ocean Night Cases BIAS



Brightness Temperature Difference 466 Clear Ocean Night Cases STD



TUNING

Microwave tuning has been discussed at NET Meetings and is in place at JPL

 $\Delta\Theta_{j,\ell}$ is tabulated for channel j, beam position ℓ and is scene independent Standard deviation of $\Theta_{j,\ell}-\Theta_{j,\ell}^c$ is constant in ℓ and comparable or less than specified channel j noise except for channel 7, which is not used

IR tuning

Computed based on 466 ocean night 9 spot clear cases based on our clear flag
Minimize differences between observed 9 spot average brightness temperature for channel i,
scene k, and brightness temperature computed using colocated ECMWF forecast

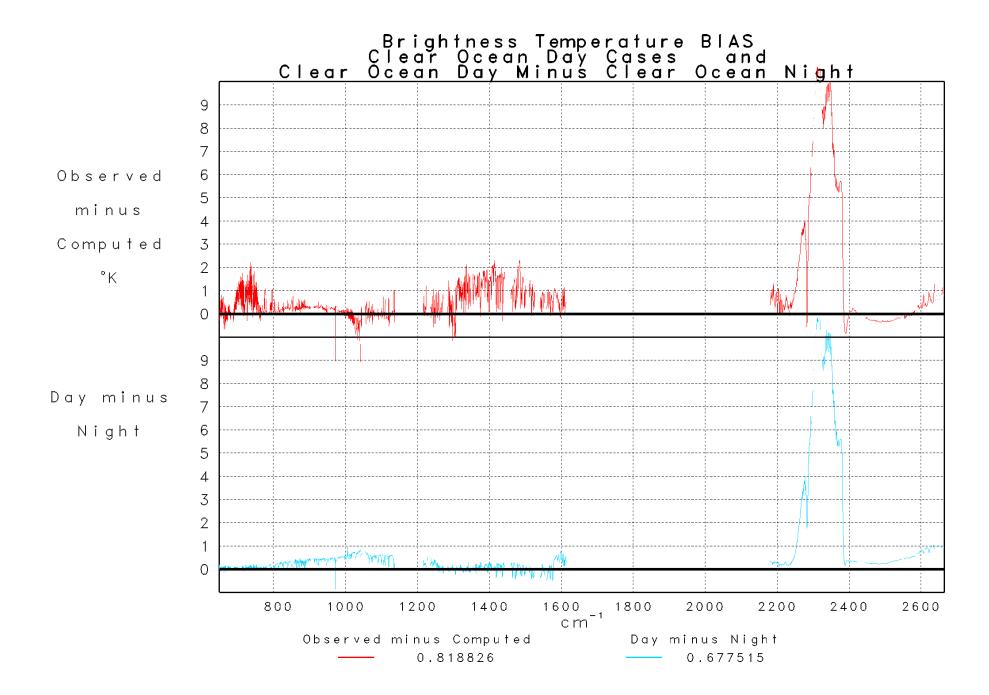
Scene independent tuning:

$$\Delta\Theta_{i,k} = A_i$$

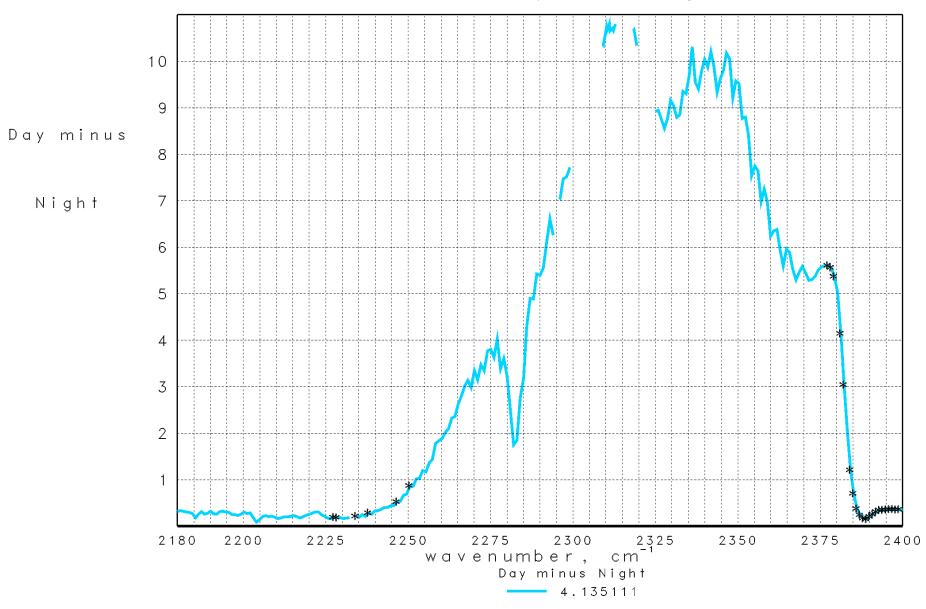
Scene dependent tuning:

 $\Delta\Theta_{i,k} = A_i + \sum B_{i,j} \left(\Theta_{j,k} - \overline{\Theta}_j\right)$ j = AMSU channels 4,5,6,8-14 Takes into account temperature profile (lapse rate) and zenith angle dependence of tuning

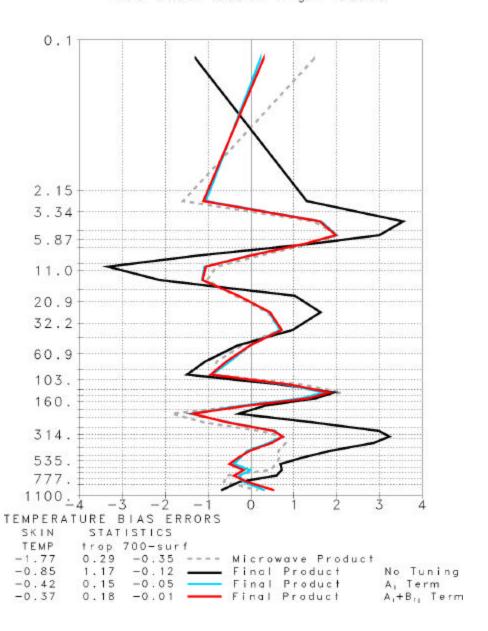
We refer to these as A_i tuning and $A_i + B_{ij}$ tuning



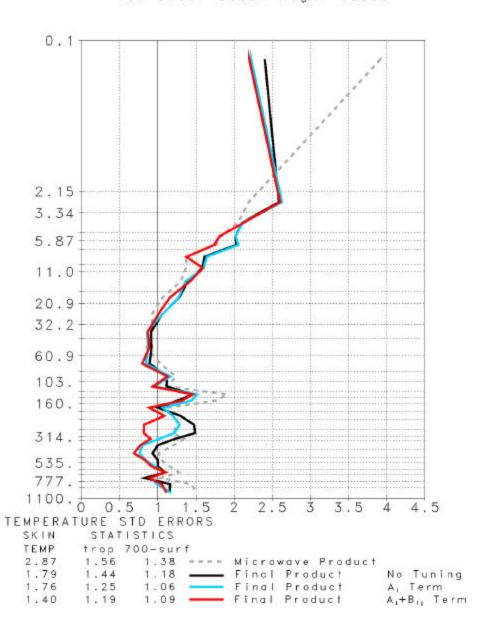
Brightness Temperature BIAS Clear Ocean Day Minus Night



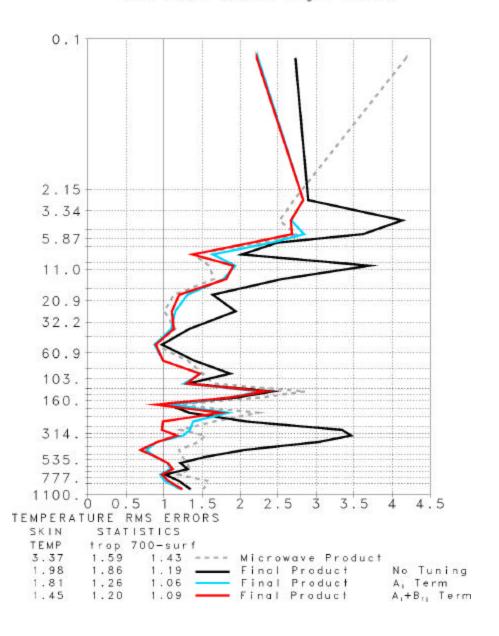
LAYER MEAN BIAS TEMPERATURE ERRORS (°C) 466 Clear Ocean Night Cases

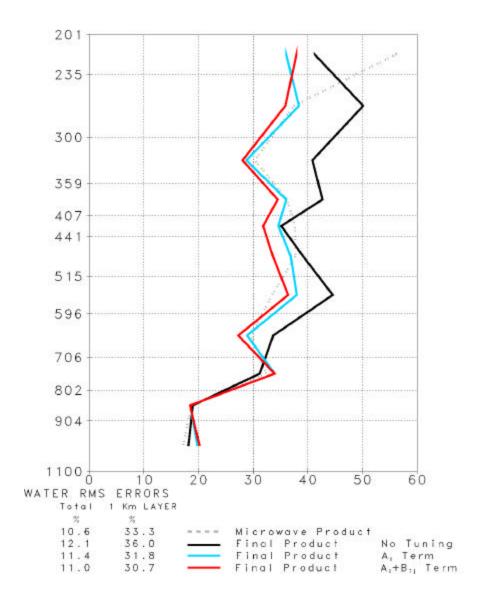


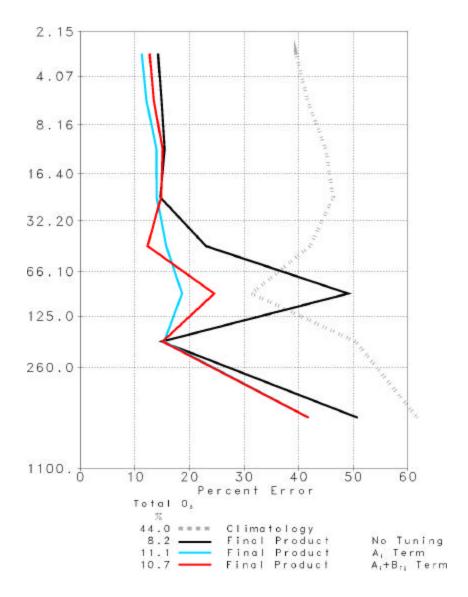
LAYER MEAN STD TEMPERATURE ERRORS (°C) 466 Clear Ocean Night Cases



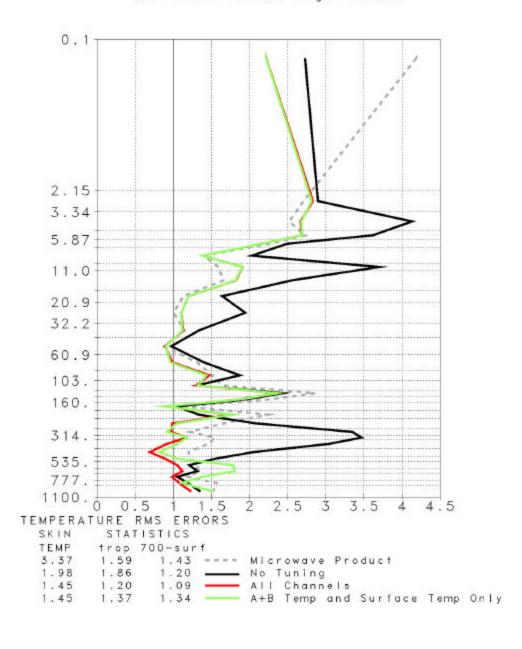
LAYER MEAN RMS TEMPERATURE ERRORS (°C) 466 Clear Ocean Night Cases



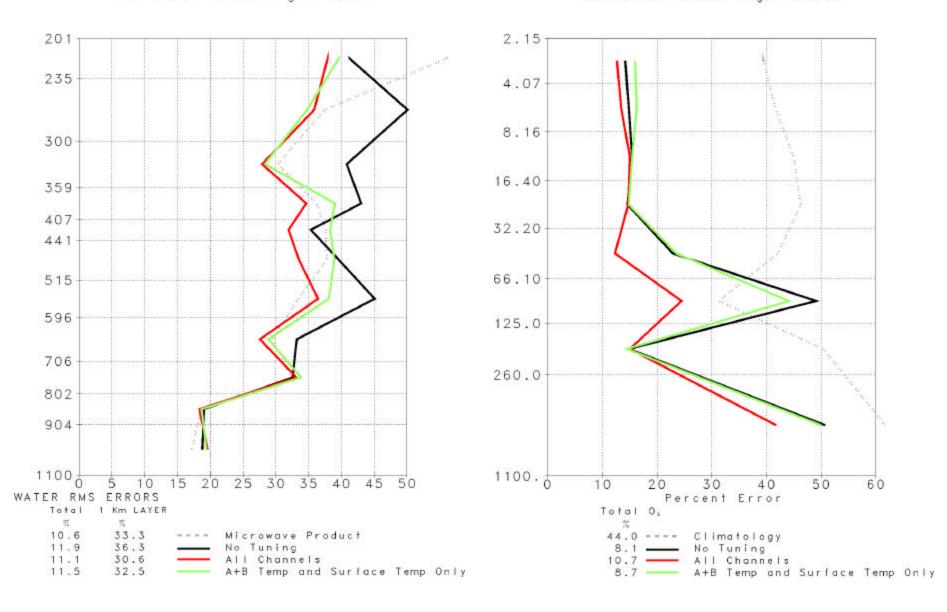




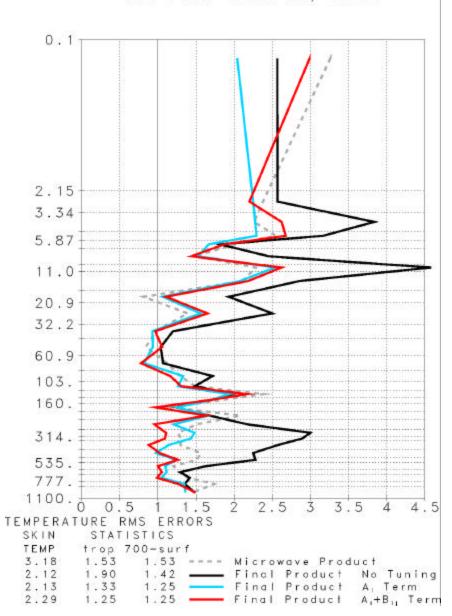
LAYER MEAN RMS TEMPERATURE ERRORS (°C) 466 Clear Ocean Night Cases

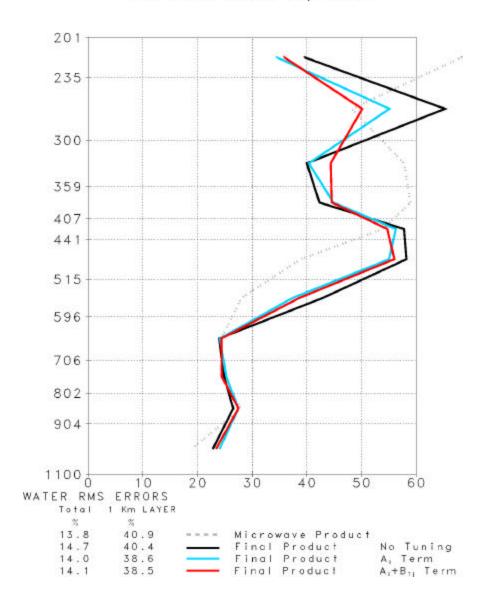


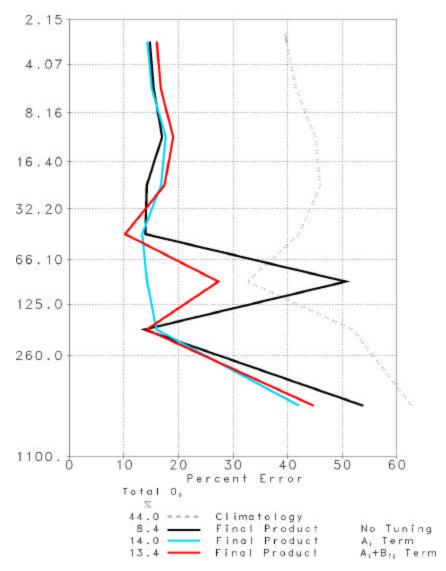
Ozone Profile RMS % Errors 466 Clear Ocean Night Cases



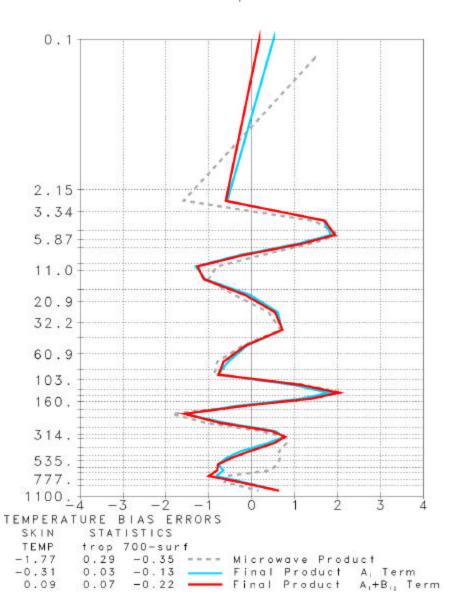
LAYER MEAN RMS TEMPERATURE ERRORS (°C) 309 Clear Ocean Day Cases



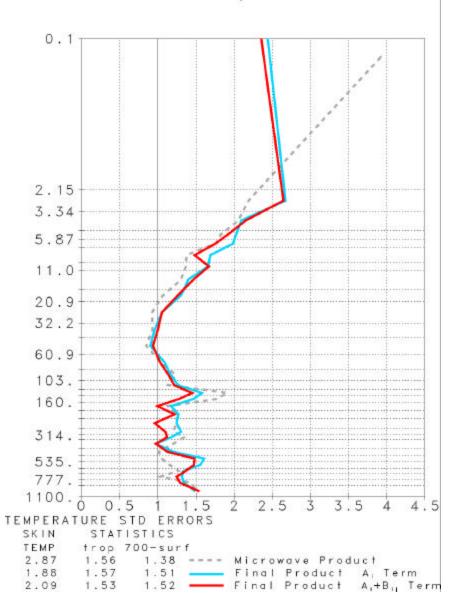




LAYER MEAN BIAS TEMPERATURE ERRORS (°C) 5116 Cloudy Ocean Cases



LAYER MEAN STD TEMPERATURE ERRORS (°C) 5116 Cloudy Ocean Cases

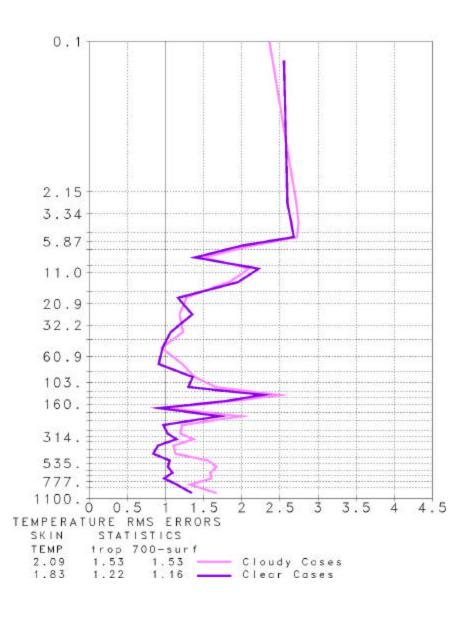


LAYER MEAN RMS TEMPERATURE ERRORS (°C) 5116 Cloudy Ocean Cases 0.1 2.15 3.34 5.87 11.0 20.9 32.2 60.9 103. 160. 314. 535. 777. 1100. 0.5 2.5 3 3.5 TEMPERATURE RMS ERRORS SKIN STATISTICS

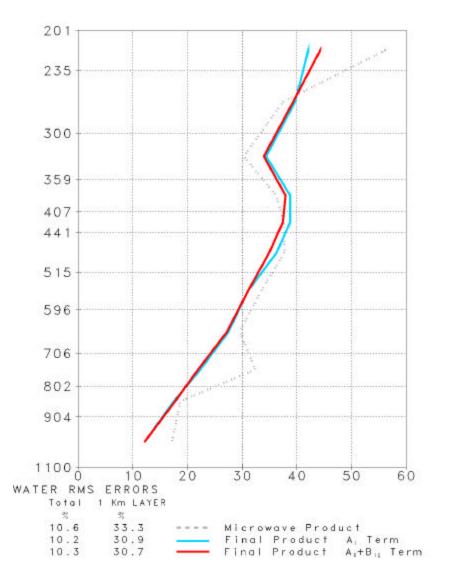
TEMP

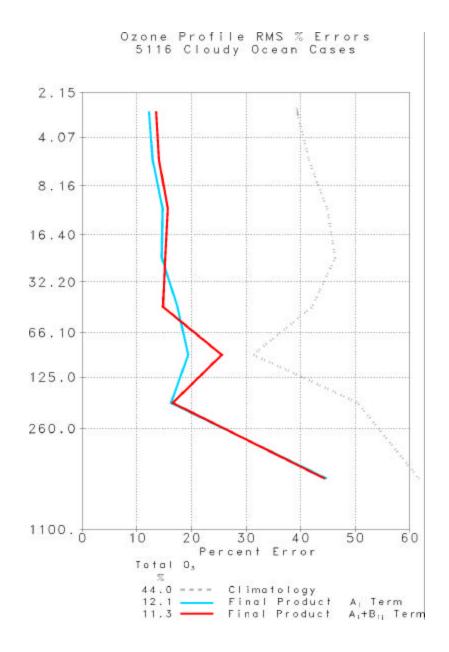
3.37 1.91 2.09

trop 700-surf

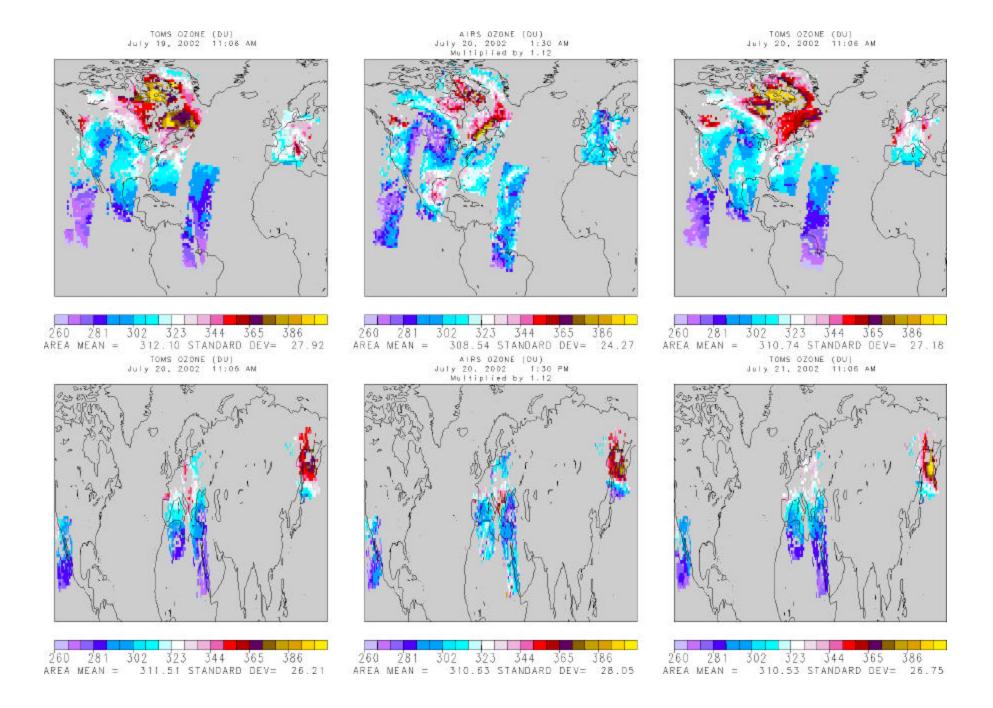


1 Km LAYER PRECIPITABLE WATER PERCENT ERRORS 5116 Cloudy Ocean Cases





Clear Column Brightness Temperature Error 5116 Cloudy Ocean Cases BIAS MITTER THAT Cloud Correction Needed °K -22 Cloud Correction Made °K -2Error °K 0 -21600 1800 2600 800 1000 1200 1400 2000 2200 2400 cm⁻ Correction Needed Correction Made Error 1.116225 **-** 1.062976 -0.053249



DIFFERENCES FROM WHAT IS CURRENTLY AT JPL

Channel noise covariance matrix at JPL does not contain estimate of computational error

Will be less important but maybe non-negligible with new RTA

IR tuning methodology is different from that of Larry McMillin running at JPL Tuning need not be iterated - it does not depend on cloud cleared radiances Constant term A_i is probably a subset of Larry McMillin's tuning matrix Matrix B_{ij} is not, unless microwave observations Θ_j are added to tuning matrix

Need for IR tuning will be considerably less once new RTA is generated A_i may be sufficient - may not even be necessary

NEAR TERM PLANS

We finally have new computer - can analyze a whole day

Once new RTA is generated

We will reassess need for IR tuning; A_i, B_{ij}

We will implement latest version of first product regression

We will assess utility of 4.3 micron channels: at night; during day

We will implement capability to run off match-up files

Compare to radiosondes

Optimize retrievals

We will continue "convergence testing" to insure GSFC and JPL code produce equivalent results